

Counting Electrons to Probe the Neutrino Mass Hierarchy

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ABSTRACT

After the successful measurement of the mixing angle θ_{13} , the determination of the neutrino mass hierarchy has become a priority for future neutrino experiments. We propose a conventional ν_μ beam with neutrino energies in the range 2-8 GeV aimed at a Mton underwater detector at the “magic” baseline of 2600 km. In this constellation it is sufficient to distinguish (ν_μ induced) track-like interactions from cascade-like interactions with moderate purity to determine the mass hierarchy.

1. Introduction

All three mixing angles and both mass square differences of the neutrino mass Eigenstates are known after the measurement of the mixing angle θ_{13} [1, 2]. Global fits of all experimental input [3] provide a coherent picture of the oscillation parameters. Among the yet unknown features of the oscillation scheme, the neutrino mass hierarchy (MH) is considered to be in reach for the next generation of experiments. It has been proposed to determine the MH by measuring atmospheric neutrinos [4] in planned low-energy extensions of existing neutrino telescopes [5, 6]. The distinction of the two hierarchy hypotheses with atmospheric neutrinos is challenging due to cancellation of contributions from neutrinos and anti-neutrinos and the effect is further attenuated by the finite energy and angular resolutions of the considered detectors. Both problems are avoided by counting beam related events of a specific flavour. Such a concept has been recently proposed [7]: muon counting from ν_μ interactions in very long baseline beams ($L > 6000$ km) should allow MH determination in Mton underwater/ice detectors. However, the construction of a steeply inclined beam-line has never been performed and is technically challenging and costly. Here, we propose instead to point a beam with a baseline of about 2600 km to a Mton underwater detector and to count ν_e interactions.

2. Oscillation Probabilities

Oscillation probabilities are calculated in a full three flavour scheme using the Globes package [8]. For a given set of neutrino parameters for normal mass hierarchy (NH), a convention is needed to transform them into inverted hierarchy (IH). We follow the approach proposed in Ref. [3], defining an average “large” mass square difference ΔM^2 which controls atmospheric neutrino oscillations

$$\Delta M^2 = \frac{1}{2} (\Delta m_{32,NH}^2 + \Delta m_{31,NH}^2) = \frac{1}{2} (\Delta m_{23,IH}^2 + \Delta m_{13,IH}^2) \quad (1)$$

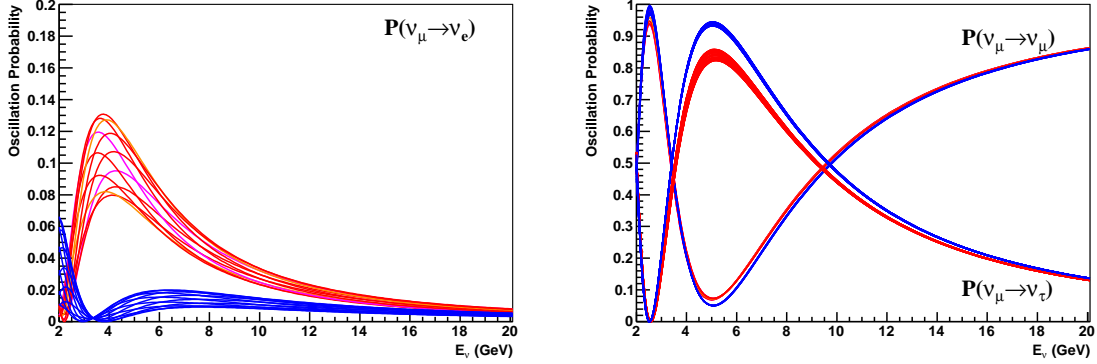


Figure 1: Neutrino oscillation probabilities $P(\nu_\mu \rightarrow \nu_e)$, $P(\nu_\mu \rightarrow \nu_\mu)$ and $P(\nu_\mu \rightarrow \nu_\tau)$ for a baseline of 2600 km and oscillation parameters from a global fit as a function of neutrino energy. The group of red lines is for NH, blue for IH, ϕ_{CP} is varied in steps of 30° between 0° and 330° . The special values $\phi_{CP} = 0^\circ, 180^\circ$ (CP conservation) are indicated in magenta, $\phi_{CP} = 90^\circ, 270^\circ$ (maximal CP violation) in orange on the left plot.

with $\Delta m_{ab}^2 = m_a^2 - m_b^2$. IH is defined as the sign change of ΔM^2 for a given set of parameters in NH. This leads to the following transformation rules for the related mass square differences:

$$\Delta m_{23,IH}^2 = \Delta m_{32,NH}^2 + \Delta m_{21}^2 \quad (2)$$

$$\Delta m_{13,IH}^2 = \Delta m_{31,NH}^2 - \Delta m_{21}^2 \quad (3)$$

Figure 1 shows the resulting oscillation probabilities $P(\nu_\mu \rightarrow \nu_e, \nu_\mu, \nu_\tau)$ with the best fit parameters from Ref. [3] for a baseline of 2600 km. Most of the neutrino path will be in the outer Earth mantle with a maximal depth of 134 km. A constant density of 3.3 g/cm^3 is used for the calculation of the oscillation probabilities as given in the PREM model [9] for the outer mantle.

For $P(\nu_\mu \rightarrow \nu_e)$ a significant difference between both hierarchies is observed. The variation of the CP-phase ϕ_{CP} (different lines of the same color) leads instead only to moderate changes of the oscillation probabilities. For $E_\nu = 3.5 \text{ GeV}$ and IH $P(\nu_\mu \rightarrow \nu_e)$ is strictly independent from ϕ_{CP} . In the range $3 \text{ GeV} < E_\nu < 8 \text{ GeV}$ there is no overlap of the two CP-bands for the two hierarchies which allows to determine MH by counting ν_e events. This special feature has been noted in the past [10] and led to the label “magic” for baselines in the range 2500-2600 km. The large value of the mixing angle $\theta_{13} \approx 9^\circ$ leads to peak values of 13% for $P(\nu_\mu \rightarrow \nu_e)$ in NH, allowing to detect a sizable sample of ν_e events in a suitable detector.

$P(\nu_\mu \rightarrow \nu_\mu)$ depends only weakly on MH and not on ϕ_{CP} . The first vacuum oscillation minimum at 5 GeV is the dominating feature in the shown energy range. Counting ν_μ events can serve as a flux normalisation or it could be used to improve

the measurement of the atmospheric oscillation parameters ΔM^2 and θ_{23} .

3. Cross Sections

The total cross sections for ν_μ and $\bar{\nu}_\mu$ charged current (CC) interactions are taken from Ref. [11] in the parton scaling approximation:

$$\sigma_{\nu_\mu}^{CC}(E_\nu) = 0.68 \cdot (E_\nu/\text{GeV}) 10^{-38} \text{cm}^2 \quad (4)$$

$$\sigma_{\bar{\nu}_\mu}^{CC}(E_\nu) = 0.34 \cdot (E_\nu/\text{GeV}) 10^{-38} \text{cm}^2. \quad (5)$$

Deviations from this linear behaviour due to quasi-elastic or resonant interactions are ignored. Their contribution would not alter the result of this study in a significant way. The relevant neutrino energies are significantly larger than m_μ and m_e therefore from flavour universality $\sigma_{\nu_e}^{CC} = \sigma_{\nu_\mu}^{CC}$ and $\sigma_{\bar{\nu}_e}^{CC} = \sigma_{\bar{\nu}_\mu}^{CC}$. However for ν_τ CC interactions the mass of the τ -lepton cannot be neglected. We use the calculation from Ref. [12]. For a threshold energy $E_0 = 5 \text{ GeV}$ and $E_0 < E_\nu < 30 \text{ GeV}$ we find the following simple parametrisation:

$$\sigma_{\nu_\tau}^{CC}(E_\nu) = 0.29 \log \left(\frac{E_\nu}{E_0} \right) \sigma_{\nu_\mu}^{CC}(E_\nu). \quad (6)$$

The cross sections for neutral current (NC) interactions for all flavours are approximated as:

$$\sigma_\nu^{NC}(E_\nu) = \frac{1}{3} \sigma_{\nu_\mu}^{CC}(E_\nu) \quad (7)$$

$$\sigma_{\bar{\nu}}^{NC}(E_\nu) = \frac{1}{3} \sigma_{\bar{\nu}_\mu}^{CC}(E_\nu). \quad (8)$$

Whereas the sum of the oscillation probabilities $\sum_\alpha P(\mu \rightarrow \alpha)$ equals unity due to the unitarity of the flavour mixing matrix, the cross section weighted sum

$$P_\mu^\sigma(E_\nu) = \frac{1}{\left[\sigma_{\nu_\mu}^{CC} + \sigma_\nu^{NC} \right] (E_0)} \sum_\alpha \left[P(\mu \rightarrow \alpha) \sigma_{\nu_\alpha}^{CC} + \sigma_\nu^{NC} \right] (E_\nu) \quad (9)$$

(normalised at $E_0 = 1 \text{ GeV}$) may help identifying the optimal energy range to separate the two MH hypotheses. P_μ^σ is shown on the left plot of Figure 2 and can be interpreted as the event rate per neutrino energy seen in a detector with an energy independent detection efficiency for a pure ν_μ flux flat in energy. No distinction is made neither between NC and CC interactions nor between different flavours. Nevertheless a clear separation of the two MH hypotheses is observed for energies above 3 GeV. It can be attributed to the kinematical suppression of ν_τ CC interactions. The size of the effect is quantified in the right plot of Figure 2 which shows the ratio between IH and NH integrals from the left plot (for different values of ϕ_{CP}) taken between an assumed

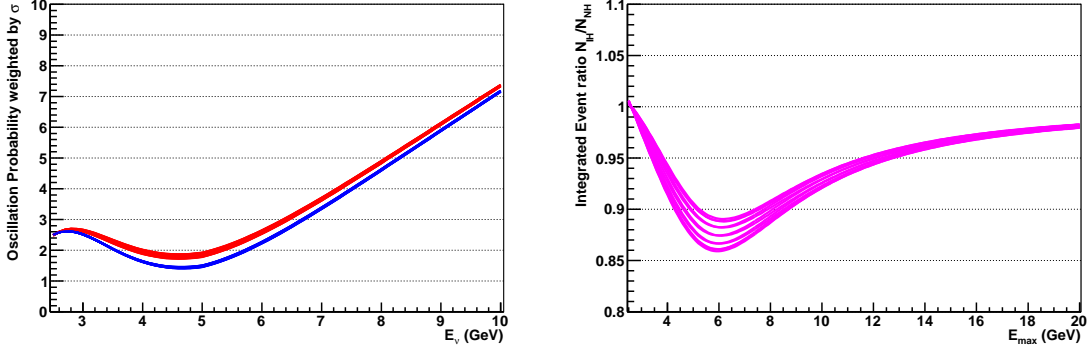


Figure 2: **Left:** Summed oscillation probabilities $P(\mu \rightarrow \alpha)$ weighted by cross sections (red : NH, blue : IH) **Right:** Ratio of Integrals N_{IH}/N_{NH} from left plot between $E_{min}=2.5$ GeV and E_{max} .

threshold energy of 2.5 GeV and a variable maximal energy E_{max} . An optimal value of $E_{max} = 6$ GeV can be read from the figure, which yields a suppression of the IH event rate of 11-14% compared to the expected rate for NH. Extending the energy range to higher values reduces the relative size of the separation of the MH hypotheses. It can be concluded that a measurement of the mass hierarchy will be possible even without any flavour tagging capabilities using a neutrino beam in a limited energy range of 2-6 GeV and a large detector which can reliably count beam related neutrino interactions. Nevertheless flavour tagging methods are discussed below and they are used to improve the significance of the measurement.

4. Neutrino beam

The Institute of High Energy Physics (IHEP) in Protvino near Moscow (located at $54^{\circ}52'N, 37^{\circ}11'E$) hosts the U70 proton accelerator [13] which provides protons with energies up to 70 GeV. It is operational since 1967 and had been the world largest proton accelerator at its time of commissioning. The beam provides $1.7 \cdot 10^{13}$ protons per pulse at a repetition rate of 0.1 Hz. A planned upgrade to a total beam power of 0.45 MW will yield $2.2 \cdot 10^{14}$ protons per pulse at a repetition rate of 0.2 Hz [13]. This will allow for the accumulation of $N_{pot} = 10^{21}$ protons on target after three years of operation assuming a running time of 6 months per year at a moderate beam efficiency of 50%. In the following, event numbers are calculated for $N_{pot} = 10^{21}$.

In the past, a neutrino beam was provided to several experiments, among them the bubble chamber SKAT [14]. Secondary hadrons were produced in an Aluminum target and focused by parabolic lenses. Neutrinos were produced in a 140 m long decay line. The bubble chamber was situated 270 m behind the target, and at a distance of $l_{SKAT} = 245$ m downstream the beginning of the decay tube. This value will be used in the following to scale the beam intensity to the remote location. The neutrino fluxes $d\Phi_{\nu}/dE_{\nu}$, as they were delivered to the SKAT experiment for focusing

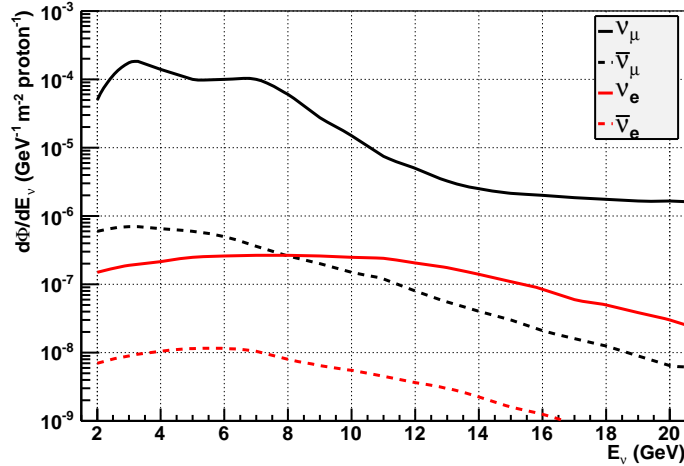


Figure 3: Neutrino Flux $d\Phi_\nu/dE_\nu$ per proton as seen by the SKAT bubble chamber. The different lines correspond to the different neutrino flavours.

of positively charged hadrons and for all four flavours present in the beam, are shown in Figure 3. A parametrisation of the beam intensity as function of the neutrino energy at the SKAT detector has been obtained from Ref. [14, 15] and will be used in the following. As seen from Figure 3, a very clean ν_μ beam was provided with energies dominantly in the range 2-8 GeV. Contaminations from other flavours were on the sub-percent level.

To provide a beam to a detector in the Mediterranean Sea at a distance of $l_{LBL} \approx 2600$ km a new beam-line is needed. It would point in the southwest direction from the proton accelerator ring, an area which is currently not obstructed by buildings. A moderate downward inclination of 11.7° is needed. A 300 m long decay tunnel could be envisaged which yields 50% more neutrinos compared to 140 m used in the past. It would end 60 m below ground. A possible near detector might be located 500 m behind the target at a depth of 100 m. The calculation of event numbers below will be based on the traditional beam [14] scaled up by a factor 1.5 to account for a decay tunnel of 300 m.

5. Detector

The ANTARES detector [16] is a deep sea neutrino telescope, which operates successfully in the Mediterranean Sea at ($42^\circ 48'N, 6^\circ 10'E$). A neutrino beam from Protvino to this location would result in a baseline of 2588 km. Recently, the ANTARES Collaboration published a measurement of atmospheric neutrino oscillations [17], demonstrating the capability of the device to detect and measure neutrinos with energies as low as 20 GeV. A multi-cubic-kilometer detector KM3NeT [18] is planned as a future neutrino telescope in the Mediterranean Sea. One possible site

would be close to the existing ANTARES detector. An alternative site in the Ionian Sea off the Sicilian coast results incidentally in an identical baseline (within 1%) for a beam from Protvino. Therefore the calculation, presented here, holds for both site options.

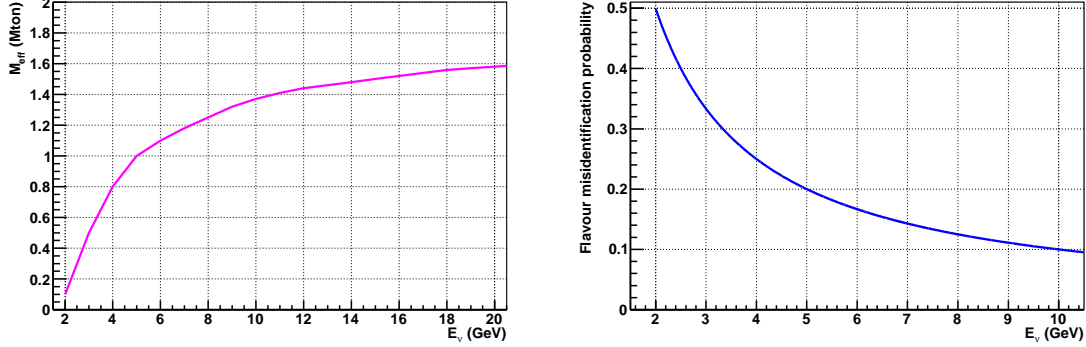


Figure 4: **Left:** Effective mass M_{eff} for the detection of ν_μ CC interactions as function of E_ν for events with the interaction vertex inside the instrumented volume. A successful track reconstruction including a condition on the likelihood of the track fit is required. **Right:** Flavour misidentification probability as function of E_ν . The same function is assumed for the probability to misidentify tracks as cascades and vice versa.

Currently the KM3NeT Collaboration carries out a feasibility study to determine the physics reach for MH determination of a densely equipped detector based on about 20% of the final budget [6]. About 1000 optical modules would be used to instrument a water mass of around 2 Mtons. The detector is supposed to be sensitive to neutrinos with energies down to few GeV. Figure 4 shows a preliminary result of this study [6]: the effective mass M_{eff} for ν_μ charged current interactions occurring inside the instrumented volume is given as function of the neutrino energy. Here a trigger condition, a likelihood based track reconstruction and a quality criteria for the goodness of the track fit are applied.

It is assumed that the considered detector is able to distinguish “track” and “cascade” event signatures. A “track” requires the presence of a long muon trajectory. As the muon range increases linearly with energy whereas the longitudinal extension of cascades remains essentially unchanged for the considered energies, the distinguishability of these two topologies will be a function of the neutrino energy. For the event rate calculation below, we introduce $\epsilon(E_\nu)$ as the probability to misidentify a cascade as track, while the misidentification of a track as cascade will be called $\eta(E_\nu)$. Both functions are parametrised by

$$\epsilon(E_\nu) = \eta(E_\nu) = 1/(E_\nu/\text{GeV}); E_\nu > 2\text{GeV} \quad (10)$$

which is illustrated in the right plot of Figure 4. For $E_\nu = 2$ GeV, $\epsilon = \eta = 0.5$, which means the two topologies cannot be distinguished and the attribution of an

event to one of them is random. A 5 GeV neutrino produces a muon with an average range of 15 m in Sea water which exceeds already by far the typical longitudinal size of a hadronic or electromagnetic shower. Correspondingly the misidentification probability is assumed to drop to 20%. For $E_\nu = 10$ GeV ϵ and η further decrease to 10%. However, these low values of the misidentification probability are not exploited in this analysis due to the energy profile of the beam, which suppresses contributions of neutrinos with $E_\nu > 10$ GeV.

6. Signal Event Rates

The rate of detected CC events of flavour α can now be calculated:

$$\frac{dN_\alpha}{dE_\nu} = N_{pot} \left(\frac{l_{SKAT}}{l_{LBL}} \right)^2 \frac{M_{eff}(E_\nu)}{m_p} \left[\sigma_{\nu_\alpha}^{CC} \left(\frac{d\Phi_{\nu_\mu}}{dE_\nu} P_{\mu\alpha} + \frac{d\Phi_{\nu_e}}{dE_\nu} P_{e\alpha} \right) + \sigma_{\bar{\nu}_\alpha}^{CC} \left(\frac{d\Phi_{\bar{\nu}_\mu}}{dE_\nu} P_{\bar{\mu}\alpha} + \frac{d\Phi_{\bar{\nu}_e}}{dE_\nu} P_{\bar{e}\alpha} \right) \right] \quad (11)$$

with m_p the proton mass and the abbreviation $P_{\beta\alpha} = P(\nu_\beta \rightarrow \nu_\alpha)$ and $P_{\bar{\beta}\alpha} = P(\bar{\nu}_\beta \rightarrow \bar{\nu}_\alpha)$ for the oscillation probabilities. All four initial-state neutrino flavours which are present in the neutrino beam are taken into account. The major contribution for all final-state flavours come from the dominant ν_μ -beam component.

Information on reconstruction efficiencies for ν_e and ν_τ CC events are not available yet. Contained ν_e and ν_μ CC events will on average deposit the same amount of energy in the detector, therefore the same neutrino energy dependent effective mass $M_{eff}(E_\nu)$ from Figure 4 is assumed for both. The same efficiency function is also used for ν_τ CC events, despite the fact that they release less energy in the detector due to the escaping neutrino(s) from the tau decay. As ν_τ interactions are a background in the present analysis, this is a conservative approximation.

Similarly the rate of detected NC events can be calculated

$$\frac{dN_{NC}}{dE_\nu} = N_{pot} \left(\frac{l_{SKAT}}{l_{LBL}} \right)^2 \frac{M_{eff}(E_\nu/2)}{m_p} \left[\sigma_\nu^{NC} \left(\frac{d\Phi_{\nu_\mu}}{dE_\nu} + \frac{d\Phi_{\nu_e}}{dE_\nu} \right) + \sigma_{\bar{\nu}}^{NC} \left(\frac{d\Phi_{\bar{\nu}_\mu}}{dE_\nu} + \frac{d\Phi_{\bar{\nu}_e}}{dE_\nu} \right) \right]. \quad (12)$$

No dependence on oscillation parameters enters here. As on average 50% of the neutrino energy is transferred to the outgoing neutrino, the effective mass M_{eff} is evaluated at $E_\nu/2$.

The “track-signal” events are then simply given by $dN_{sig}^{track}/dE_\nu = dN_\mu/dE_\nu$, whereas the “cascade-signal” will be defined as $dN_{sig}^{casc}/dE_\nu = dN_e/dE_\nu$. ν_τ CC and NC interactions are considered as background. Figure 5 shows the expected signal event rates. The corresponding integral event numbers are summarized in Table 1. At this stage we do not consider the misidentification between the two event topologies,

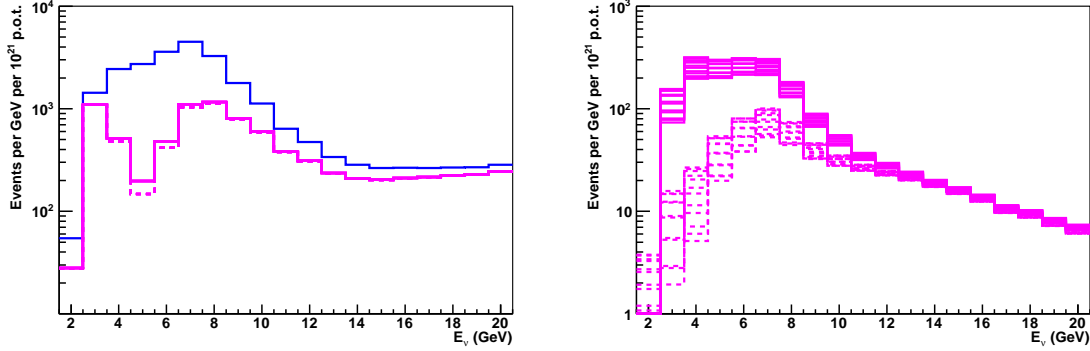


Figure 5: Observed events for $N_{pot} = 10^{21}$ as function of the neutrino energy. **Left:** Track-like events from ν_μ CC interactions. **Right:** Cascade-Like events from ν_e CC interactions. The solid blue histogram on the left plot is obtained by ignoring neutrino oscillations. Solid magenta histograms are for NH and various values of ϕ_{CP} , dashed lines indicate IH.

i.e. $\epsilon = \eta = 0$. The ν_μ CC rate is suppressed by 60% due to neutrino oscillations compared to the no-oscillation hypothesis. Nonetheless a comfortable event sample of 10000 events can be detected for $N_{pot} = 10^{21}$ with a statistical uncertainty of 1% whereas the two mass hierarchy hypotheses modify the expected event rate by 3%.

Channel	Tracks NH	Tracks IH	Cascades NH	Cascades IH
No oscil	26315		—	
Signal	10317	10015	1366-1876	397-597
ν_τ	227-231	245-248	1076-1098	1163-1176
NC	0	0	4732	4732
BG Total	227-231	245-248	5807-5830	5895-5908
Total	10543-10548	10260-10263	7196-7683	6304-6492

Table 1: Event numbers for $N_{pot} = 10^{21}$ in the track and cascade channel for both mass hierarchy schemes and varying ϕ_{CP} values with a perfect separation of track and cascade signatures.

The right plot of Figure 5 shows the event rate of ν_e CC events. For NH 1621 ± 255 events are expected, where the uncertainty is due to the unknown CP-phase. This has to be compared to 497 ± 100 events for IH. The statistical separation of both samples is better than 20σ . The largest effect is seen for neutrino energies from 3-8 GeV, as expected. Despite the fact that Figure 5 illustrates the E_ν distribution of the selected event samples, no assumption about the energy determination is needed here to determine MH. However, a moderate energy resolution will certainly increase the significance of the MH hypothesis test.

7. Background and Purity of Event Selection

The measurement will be complicated by background. Contributions from at-

mospheric neutrinos and misreconstructed down-going atmospheric muons can be ignored. A pulsed beam with a typical duty cycle lower than 10^{-6} allows to safely discard these events. The beam itself is a source of background events. ν_τ CC events with a muonic τ -decay produce “track” events, whereas NC events and ν_τ CC events with a non-muonic τ -decay have a genuine cascade signature. These two contributions are added in Table 1 and they are the only background contributions, if we assume a perfect flavour tagging mechanism to separate track-like from cascade-like events. Whereas the sample of track-like events is only marginally affected by the addition of ν_τ events, the cascade event sample is now dominated by NC events which contribute about three times as much as the ν_e signal events.

In a real detector it will not be possible to separate track-like and cascade-like events with 100% efficiency as discussed in Section 4. The probability to misidentify a cascade as track $\epsilon(E_\nu)$ and the misidentification probability of a track as cascade $\eta(E_\nu)$ as introduced in Section 5 will be used now. With these two quantities the total background for the two event samples can be written

$$\begin{aligned}\frac{dN_{bg}^{track}}{dE_\nu} &= \epsilon \frac{dN_{sig}^{casc}}{dE_\nu} + [\epsilon(1 - BR_{\tau\mu}) + (1 - \eta)BR_{\tau\mu}] \frac{dN_\tau}{dE_\nu} + \epsilon \frac{dN_{NC}}{dE_\nu} \\ \frac{dN_{bg}^{casc}}{dE_\nu} &= \eta \frac{dN_{sig}^{track}}{dE_\nu} + [(1 - \epsilon)(1 - BR_{\tau\mu}) + \eta BR_{\tau\mu}] \frac{dN_\tau}{dE_\nu} + (1 - \epsilon) \frac{dN_{NC}}{dE_\nu}.\end{aligned}\tag{13}$$

$BR_{\tau\mu}$ stands for the muonic branching ratio of the tau decay (17.4%). The total number of observed events in each channel is given by adding the (reduced) number of signal events and the background contribution from Equation 13:

$$\begin{aligned}\frac{dN_{tot}^{track}}{dE_\nu} &= (1 - \eta) \frac{dN_{sig}^{track}}{dE_\nu} + \frac{dN_{bg}^{track}}{dE_\nu} \\ \frac{dN_{tot}^{casc}}{dE_\nu} &= (1 - \epsilon) \frac{dN_{sig}^{casc}}{dE_\nu} + \frac{dN_{bg}^{casc}}{dE_\nu}.\end{aligned}\tag{14}$$

The resulting rates are shown in Figure 6 and quoted in Table 2. The signal contribu-

Channel	Tracks NH	Tracks IH	Cascades NH	Cascades IH
No oscil	26315		—	
Signal	8990	8735	1134-1547	350-519
Misreco	232-329	47-79	1326	1280
ν_τ	324-332	351-355	978-998	1057-1068
NC	1092	1092	3640	3640
BG Total	1655-1745	1494-1522	5944-5964	5977-5988
Total	10645-10736	10229-10257	7099-7491	6338-6496

Table 2: Event numbers for $N_{pot} = 10^{21}$ in the track and cascade channel for both mass hierarchy schemes and varying ϕ_{CP} values.

tions are reduced by 15-20%. The track-like sample is still dominated by signal. The different background channels add up to 15-18% of the overall rate. The situation for the cascade-like events does not change very much compared to the numbers given in Table 1. Some backgrounds are reduced due to migration into the track-like channel (e.g. NC), others are increased. The total number of background events is almost unaffected by the choice of MH (and by ϕ_{CP}) as the MH dependence of different backgrounds has the tendency to cancel. This leaves the event difference of the signal part almost unaffected. The event rates between the two MH hypotheses differ now by 9-18% with a statistical uncertainty of 1.2%. The statistical significance of the MH hypothesis test is still better than 7σ and it remains at the level of 3σ even when adding an additional systematic uncertainty of 3-4% (depending on the true value of the CP phase) for the determination of the total cascade event rate. The knowledge of the detector performance, water parameters, neutrino cross sections, oscillation parameters and the monitoring of the neutrino flux contribute to the systematic uncertainty. The neutrino flux normalisation can be controlled by performing a complementary measurement of ν_μ CC events. Uncertainties of the oscillation parameters will have been reduced by ongoing experiments and neutrino cross sections will have been measured with high precision by ongoing and planned short-baseline experiments by the time the here proposed experiment is running. The water in the abyss of the Mediterranean Sea is extensively studied in the ANTARES experiment. All these measurements will help reducing the systematic uncertainty.

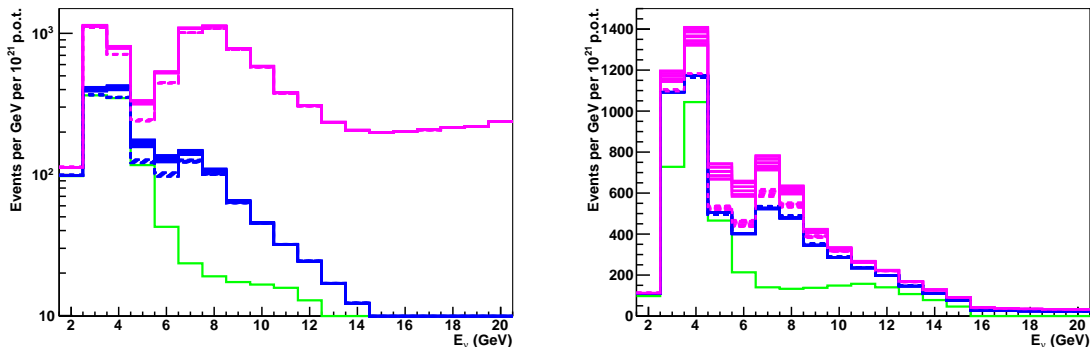


Figure 6: Observed events for $N_{pot} = 10^{21}$ as function of the neutrino energy. **Left:** Track-like events. **Right:** Cascade-Like events. In green the background from NC events is given. The blue histogram shows the total background including misreconstructed CC events. Magenta histograms are for the total event rate (signal plus background) for NH (solid) and IH (dashed) and various values of ϕ_{CP} .

Figure 6 illustrates the total event numbers as function of neutrino energy, detailing the different contributions. NC events are added to the $E_\nu/2$ bins according to their lower light yield in the detector. Their main contribution is found for energies below 6 GeV. Background from CC events is instead mainly seen above 5 GeV. The

separation of the MH hypotheses is most pronounced in the range 4-8 GeV.

8. Conclusion

A neutrino beam from IHEP Protvino to a Mton detector installed in the abyss of the Mediterranean Sea has a baseline close to the “magic” value of 2600 km. A low energy “phase-1” part of the future KM3NeT neutrino telescope could serve as target for such a beam. Counting of cascade-like events would allow a measurement of the neutrino mass hierarchy with a significance of 3σ for 10^{21} protons on target and a systematic uncertainty of 3-4% for the event rate determination. Higher values of the significance can be reached by reducing the systematic uncertainty, considering a moderate capability to measure the neutrino energy of the signal events or an improved flavour tagging method. The proposed measurement would be complementary to an analysis of atmospheric neutrinos in the same detector. The combination of both measurements will possibly allow an unambiguous determination of the neutrino mass hierarchy.

9. Acknowledgment

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10. References

- 1) F.P. An *et al.* [Daya-Bay Collab.], Phys. Rev. Lett. **108** (2012) 171803.
- 2) J.K. Ahn *et al.* [Reno Collab.], Phys. Rev. Lett. **108** (2012) 191802.
- 3) G.L. Fogli *et al.*, arXiv:1205.5254v3.
- 4) E. Kh. Akhmedov, S. Razzaque and A. Yu. Smirnov, arXiv:1205.7071v5.
- 5) J. Koskinen, presentation at “New Directions in Neutrino Physics”, Aspen, February 2013, <https://indico.cern.ch/conferenceDisplay.py?confId=224351>.
- 6) J. Brunner, presentation at “New Directions in Neutrino Physics”, Aspen, February 2013, <https://indico.cern.ch/conferenceDisplay.py?confId=224351>.
- 7) C. Lujan-Peschard, G. Pagliaroli and F. Vissani, arXiv:1301.4577v1.
- 8) P. Huber, M. Lindner and W. Winter, arXiv:hep-ph/0407333, P. Huber *et al.*, arXiv:hep-ph/0701187.
- 9) A.M. Dziewonski and D.L. Anderson, Phys. Earth Plan. Int. **25** (1981) 297.
- 10) S.K. Raut, R. Shanker Singh and S. Uma Sankar, arXiv:0908.3741v4.
- 11) J. Behringer *et al.* [Particle Data Group], Phys. Rev. **D86** (2012) 010001.
- 12) Y.S. Jeong and M.H. Reno, arXiv:1007.1996.

- 13) Y. Kudenko, presentation at “Laguna-LBNO General Meeting”, Hamburg February 2013, <http://laguna.ethz.ch/indico/conferenceDisplay.py?confId=8>.
- 14) V.V. Ammosov *et al.* [SKAT Collab.], Z. Phys. **C40** (1988) 487.
- 15) R. Nahnauer, DESY, private communication.
- 16) M. Ageron *et al.* [ANTARES Collab.], Nucl. Instrum. Meth. **A656** (2011) 11.
- 17) S. Adrian-Martinez *et al.* [ANTARES Collab.], Phys. Lett. **B714** (2012) 224.
- 18) KM3NeT Technical Design Report (2010), ISBN 978-90-6488-033-9, available from <http://www.km3net.org>.